

Freeze-out hypersurface and resonance decay influence on femtoscopic observables

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Outline

- **Single freeze-out approximation**
 - Theoretical foundations
 - Numerical implementation
 - Freeze-out hypersurface
- **Femtoscopy definitions**
 - Separation distribution
 - Correlation function
 - Gaussian fit
 - The meaning of “HBT radii”
- **Results from Therminator**
 - Identical particles
 - Separation distributions
 - Correlation functions
 - Influence of the strongly-decaying resonances
 - Influence on size
 - Influence on λ
 - Influence on the shape of the correlation function
 - Non-identical correlations

Single freeze-out approximation analytic version (the Cracow Model)

- Developed by W. Broniowski and W. Florkowski
- Main features and assumptions of the model
 - Matter is thermally equilibrated. The abundances of particles are determined by the temperature, through statistical model. The temperature at which particles are emitted is the same ($T_{chem} = T_{kin} = T$)
 - Complete treatment of resonances (381 particle types)
 - Special choice of freeze-out hypersurface: ($\tau = \sqrt{t^2 - x^2 - y^2 - z^2}$)
 - Only 4 parameters T, μ_B (fixed by abundance ratios), proper time at freeze-out τ and transverse size ρ_{max} .
 - Hubble-like flow: $u^\mu = x_\mu / \tau = t/\tau (1, x/t, y/t, z/t)$

Numerical implementation: THERMINATOR

- THERMINATOR starts from a full emission function of primordial particles (no resonance feed-down yet):

$$S(x, p) = \frac{dN}{dp_T dy d\phi_p d\alpha_T d\alpha_{\parallel} d\phi_s}$$

Comp. Phys. Comm.
174 (2006) 669-687

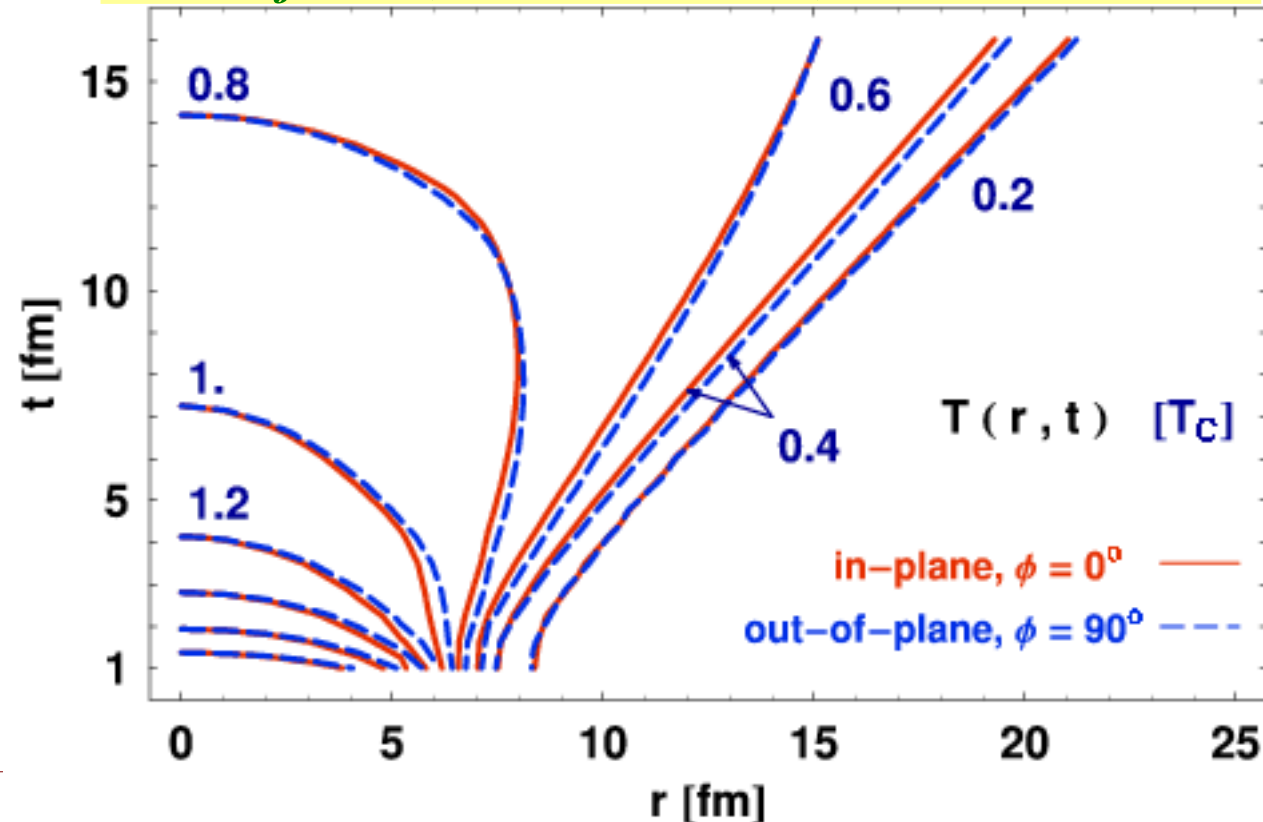
- It performs a Monte-Carlo integration of the emission function in six dimensions and then proceeds to decay all unstable particles, also with a Monte-Carlo method
 - Gives space-time freeze-out points for all particles
 - Emission function based on Blast-wave – has strong flows
 - Relative and absolute multiplicities taken from the chemical model, with parameters fit to the data

nucl-th/0504047 <http://hirc.if.pw.edu.pl/en/therminator/>

The importance of freeze-out hypersurface

- The Cracow single freeze-out model implemented a particular shape of freeze-out hypersurface where $\tau = \text{const}$
- Commonly used Blast-wave models have a hypersurface defined as $t = \text{const}$
- Hydrodynamic calculations usually produce a different shape in the t - ρ plane, so “generalized Blast-wave” was used

M. Chojnacki, W. Florkowski, nucl-th/0603035



Different freezeout hypersurface – “BlastWave” with resonances

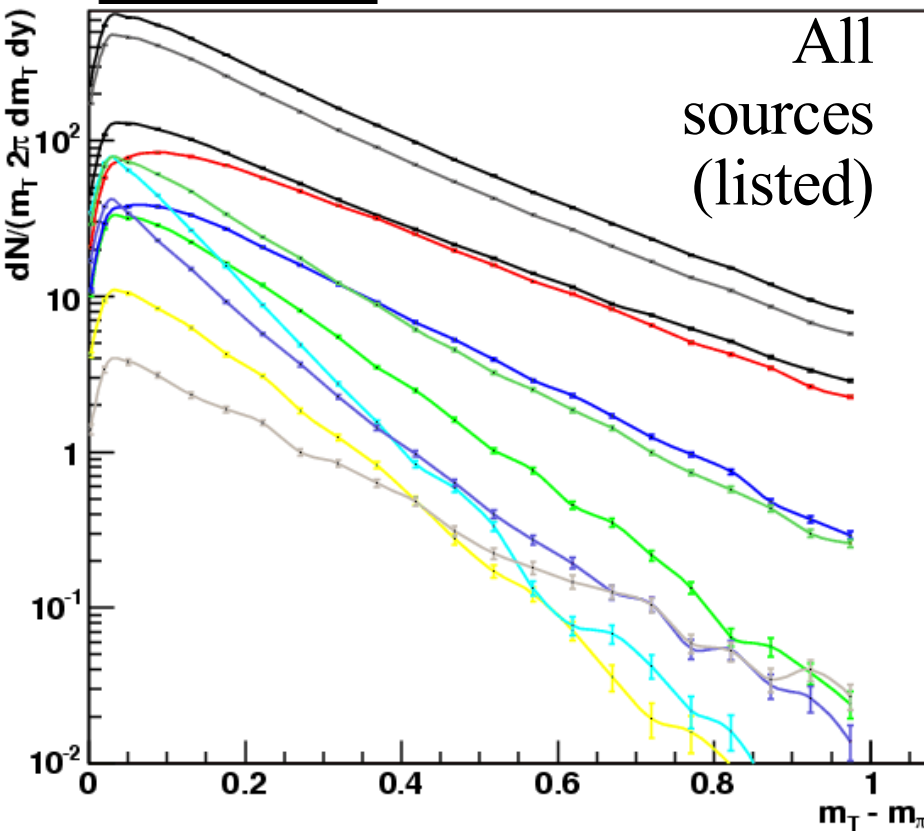
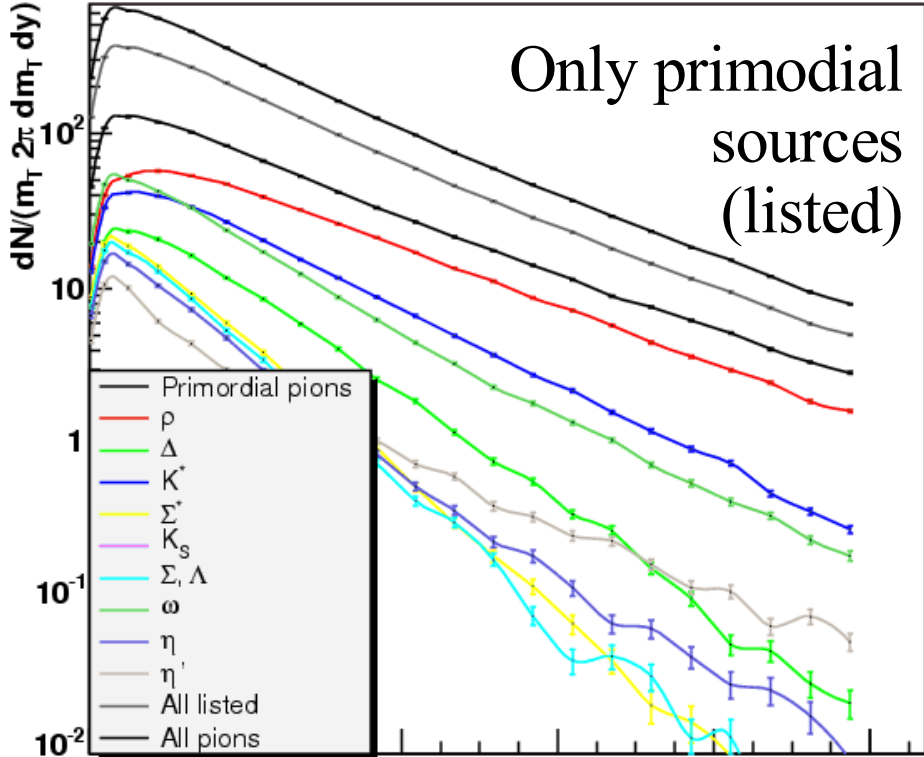
- In Therminator we have complete freedom of choice of the emission function. We use e.g. “generalized BlastWave”:

$$\frac{dN}{dp_{\perp} d\phi_P dy d\rho d\phi_S d\alpha_{\parallel}} = \frac{p_T \rho}{(2\pi)^3} (\tau + a\rho) [m_{\perp} \cosh(\alpha_{\parallel} - y) - a p_{\perp} \cos(\phi_P - \phi_S)]$$

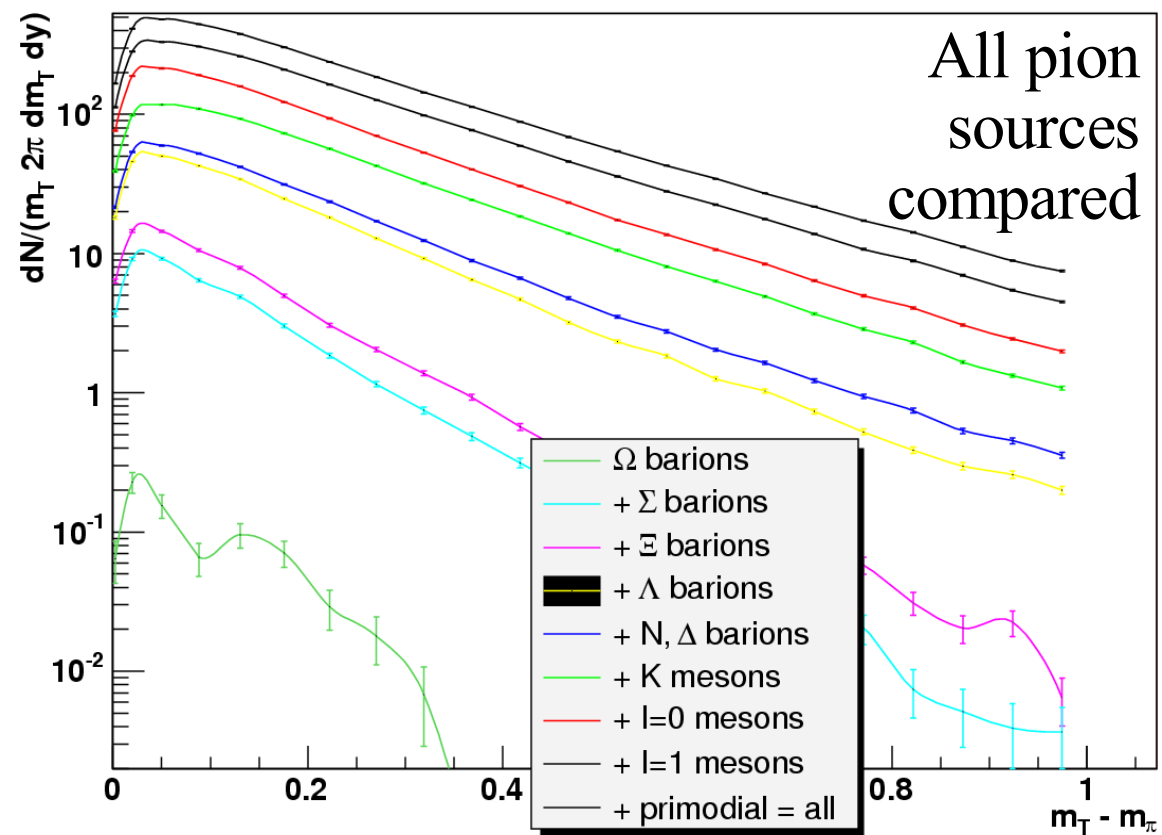
$$\times \left\{ \exp \left[\beta m_{\perp} \frac{1}{\sqrt{1 - v_{\perp}^2}} \cosh(\alpha_{\parallel} - y) - \beta p_{\perp} \frac{v_{\perp}}{\sqrt{1 - v_{\perp}^2}} \cos(\phi_P - \phi_S) - \beta \mu \right] \pm 1 \right\}^{-1}$$

- The a parameter from the $\tau + a\rho$ term determines the shape of the freeze-out in ρ - t plane
- Thermodynamical parameters (T, μ_B) stay the same. ρ_{max} and τ and v_T fitted, so that STAR pion and kaon spectra is reproduced.
- Please note this “Blast-wave” has all resonances included

Resonance influence on (pion) slopes



- Please note that "most important" resonances account for 60%(73%) of all pions (depending on how you count)



Emission function

- We use the “emission function” $S(x,p)$; a probability to emit particle with momentum p from point x .
- We can then define single particle spectra:

$$P_1(\vec{p}) = E \frac{dN}{d^3 p} = \int d^4 x S(x, p)$$

as well as two-particle spectra:

$$P_2(\vec{p}_a, \vec{p}_b) = E_a E_b \frac{dN}{d^3 p_a d^3 p_b} = \int S(\mathbf{x}_1, \mathbf{x}_2, \mathbf{p}_1, \mathbf{p}_2) d^4 x_1 d^4 x_2$$

- And the correlation function is:

$$C(\vec{q}, \vec{K}) = \frac{P_2^C(\vec{p}_a, \vec{p}_b) \delta(\vec{q} - \vec{p}_a + \vec{p}_b) \delta(\vec{K} - 1/2(\vec{p}_a + \vec{p}_b))}{P_2(\vec{p}_a, \vec{p}_b) \delta(\vec{q} - \vec{p}_a + \vec{p}_b) \delta(\vec{K} - 1/2(\vec{p}_a + \vec{p}_b))}$$

Connection to separation distribution

- If we neglect the event-wide correlations and leave only femtoscopical (two-particle) ones, we write:

$$\int S(\mathbf{x}_1, \mathbf{x}_2, \mathbf{p}_1, \mathbf{p}_2) d^4 \mathbf{x}_1 d^4 \mathbf{x}_2 \approx \int S_1(\mathbf{x}_1, \mathbf{p}_1) S_2(\mathbf{x}_2, \mathbf{p}_2) d^4 \mathbf{x}_1 d^4 \mathbf{x}_2$$

- And we can define the separation distribution

$$S(\mathbf{r}^*, p_a, p_b) \stackrel{\text{def}}{=} \int S_1(x_1, p_a) S_2(x_2, p_b) \delta(r^* - \mathbf{x}_1 + \mathbf{x}_2) d^4 x_1 d^4 x_2 = \int S_1(x_1, p_a) S_2(r^* - x_1, p_b) d^4 x_1$$

- In the uncorrelated case we have:

$$P_2(\vec{p}_a, \vec{p}_b) = \int S(\mathbf{x}_1, \mathbf{x}_2, \mathbf{p}_1, \mathbf{p}_2) d^4 x_1 d^4 x_2 = \int S(\mathbf{r}^*, p_a, p_b) d^4 r$$

- While in the correlated case we need the separation distribution itself:

$$P_2^C(p_a, p_b) = \int S(\mathbf{r}^*, p_a, p_b) |\Psi(\mathbf{r}^*, \mathbf{k}^*)|^2 d^4 r$$

The wave function

- For the correlation function we need the wave function
- For identical non-charged particles we have:

$$|\Psi(\mathbf{r}^*, \mathbf{k}^*)|^2 = 1 + \cos(\mathbf{q}_{inv} \cdot \mathbf{r}^*)$$

- If the particles are charged, then:

$$\Psi(\mathbf{k}^*, \mathbf{r}^*) = e^{i\delta_c} \sqrt{A_c(\eta)} \frac{1}{\sqrt{2}} \left[e^{-i\mathbf{k}^* \cdot \mathbf{r}^*} F(-i\eta, 1, i\xi^+) \pm (-1)^S e^{i\mathbf{k}^* \cdot \mathbf{r}^*} F(-i\eta, 1, i\xi^-) \right]$$

- If they are nonidentical, the correlation function can also be calculated. We have then:

$$\Psi(\mathbf{k}^*, \mathbf{r}^*) = e^{i\delta_c} \sqrt{A_c(\eta)} \left[e^{-i\mathbf{k}^* \cdot \mathbf{r}^*} F(-i\eta, 1, i\xi) \right]$$

where $\xi^{\pm} = \mathbf{k}^* \cdot \mathbf{r}^* \pm k^* r^* \equiv \rho(1 \pm \cos(\theta^*))$, $\rho = k^* r^*$, $\eta = (k^* a)^{-1}$, $a = (\mu z_1 z_2 e^2)^{-1}$

$A_c(\eta)$ is the Gamow factor, and F is the confluent hypergeometric function.

Numerical methods

- All analytic methods can be naturally transformed into numerical ones by performing “hadronization” using a modification of Monte-Carlo integration of the Cooper-Frye formula. Generating sets of random numbers (x,p) distributed according to the probability density $S(x,p)$, we can interpret each of them as a particle and combine them in pairs.
- Then, by the definition of the Monte-Carlo procedure, each integral of the emission function can be turned into a simple pair counting (with or without weight). E.g.
$$\int d^4 r S(\mathbf{r}^*, \mathbf{p}_1, \mathbf{p}_2) \stackrel{\text{def}}{=} \sum_i \sum_{j \neq i} \delta(\mathbf{p}_i - \mathbf{p}_a) \delta(\mathbf{p}_j - \mathbf{p}_b),$$
$$\int d^4 r S(\mathbf{r}_*, \mathbf{p}_1, \mathbf{p}_2) |\Psi|^2 \stackrel{\text{def}}{=} \sum_i \sum_{j \neq i} \delta(\mathbf{p}_i - \mathbf{p}_a) \delta(\mathbf{p}_j - \mathbf{p}_b) |\Psi|^2$$
- This is the procedure used in our studies

Radius extraction in data

- In model studies the known fit form for quantum-statistics only correlation function is used, giving “HBT radii”:

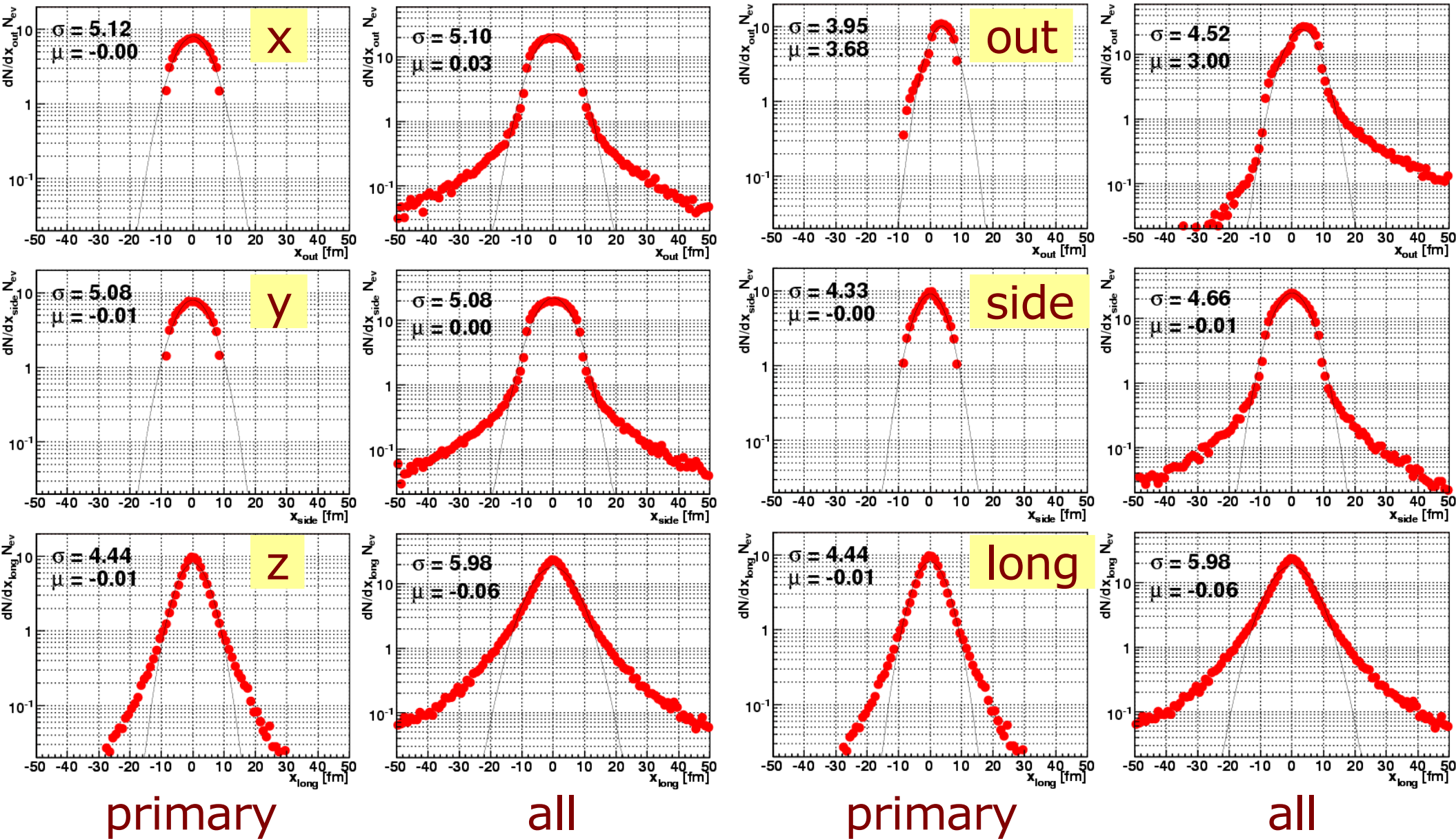
$$C(\vec{q}) = 1 + \lambda \exp(-R_{out}^2 q_{out}^2 - R_{side}^2 q_{side}^2 - R_{long}^2 q_{long}^2)$$

- To fit the experimental data or the QS+Coulomb model function, an approximation is used stating that Coulomb and symmetrization of the wave function factorize, which gives “Bowler-Sinyukov” formula:

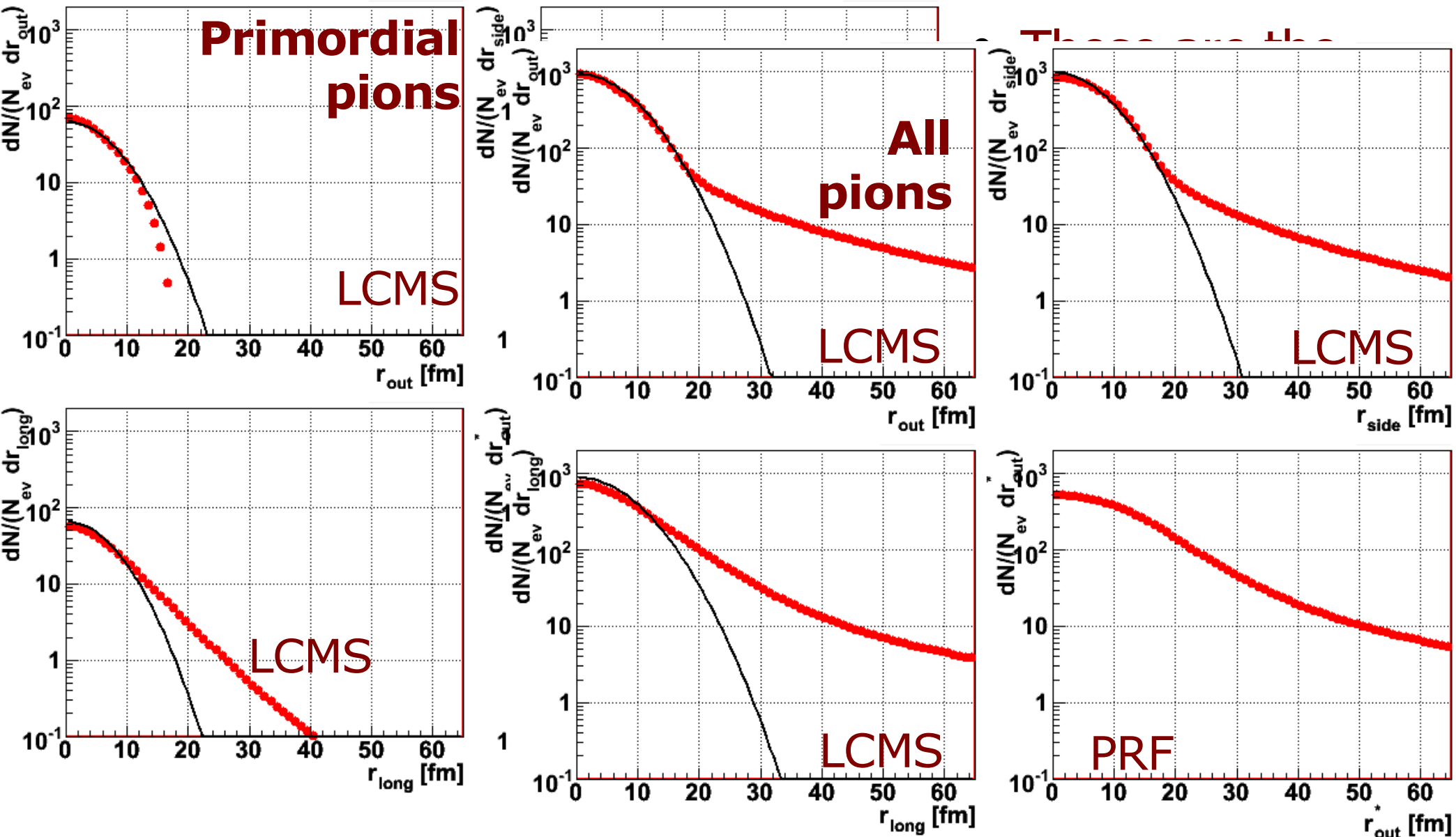
$$C(\vec{q}) = (1 - \lambda) + \lambda K_{coul}(q_{inv}) (1 + \exp(-R_{out}^2 q_{out}^2 - R_{side}^2 q_{side}^2 - R_{long}^2 q_{long}^2))$$

- K_{coul} is the Coulomb-only wave-function averaged over a 3D gaussian with the same, fixed size in all directions, .
- Following the experiment all analysis is done in LCMS

Single particle distributions

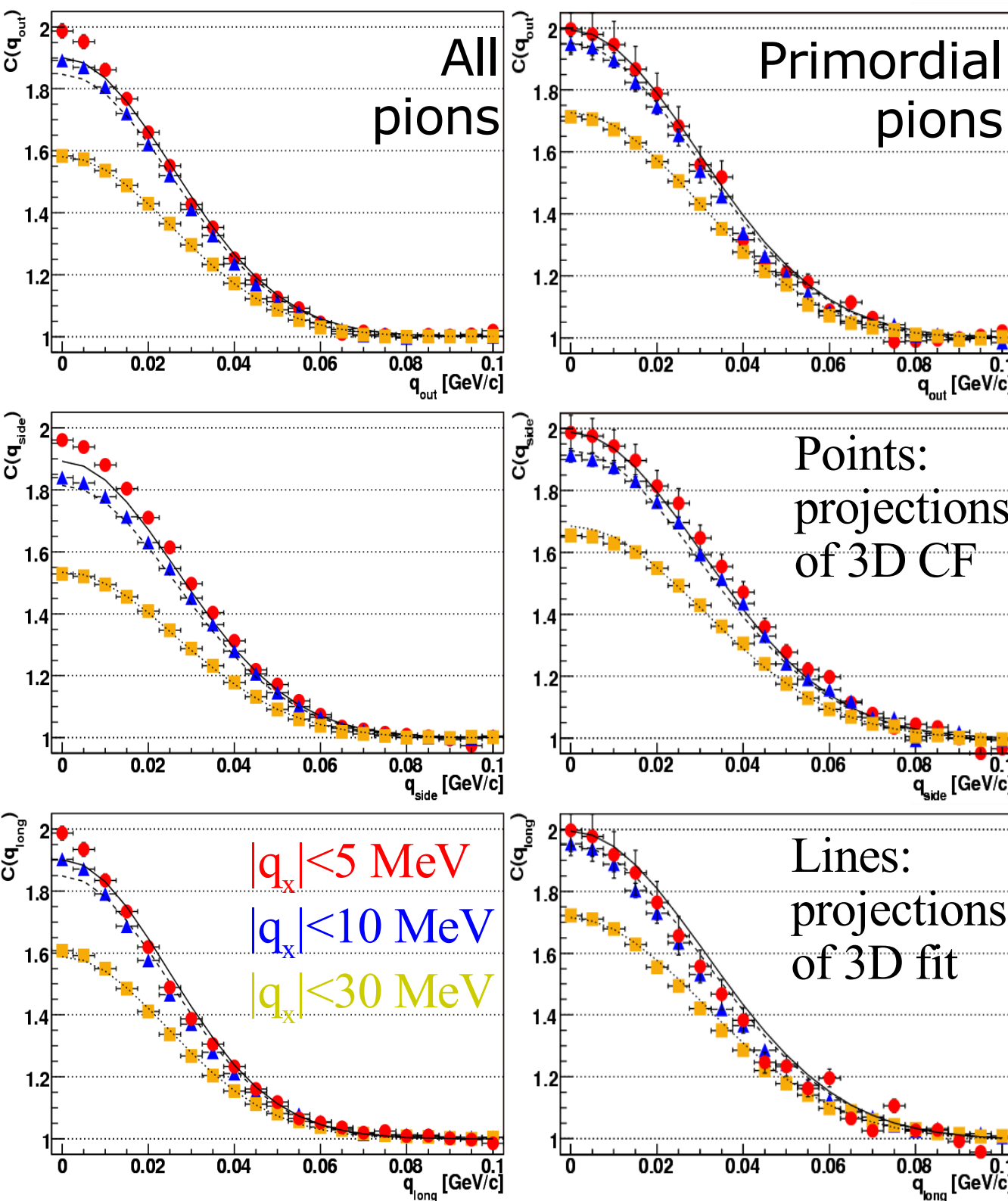


Two-particle separation distributions

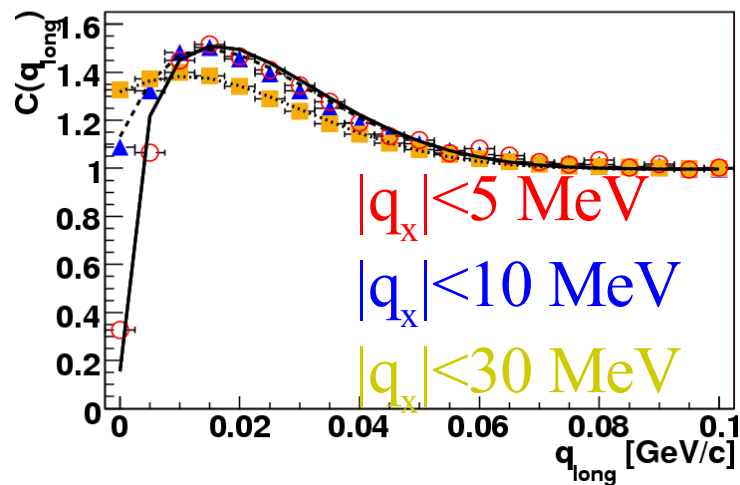
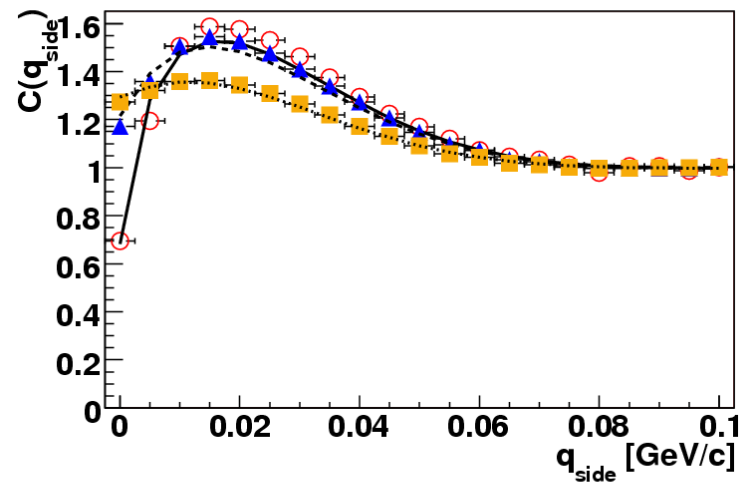
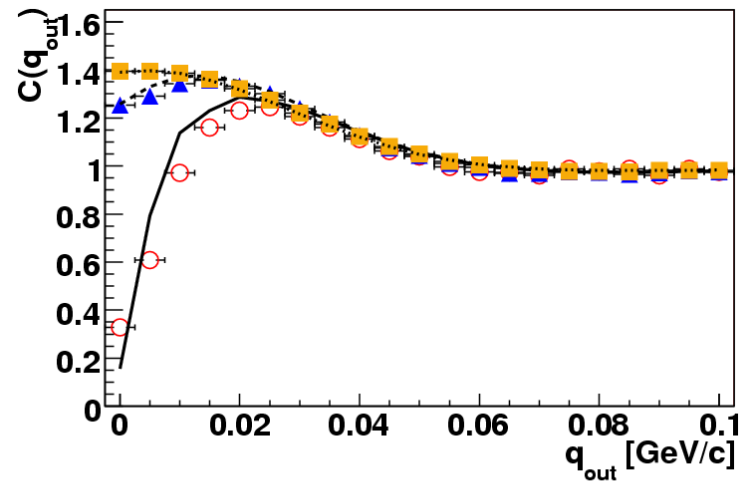


3D correlation functions

- We calculate 3D correlation functions and fit them with a traditional formula
- The fits perform relatively well, but the functions are significantly non gaussian (as expected!)

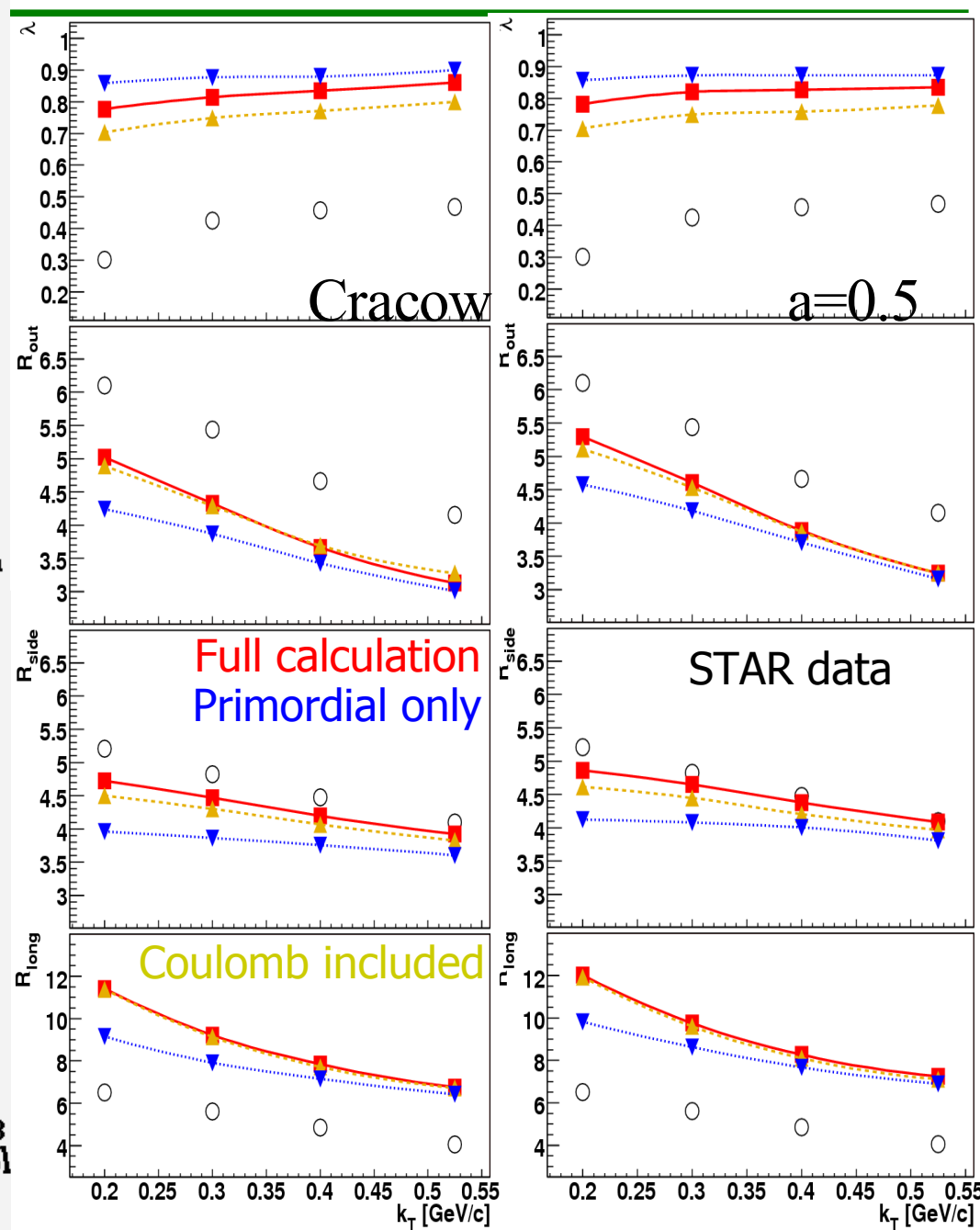
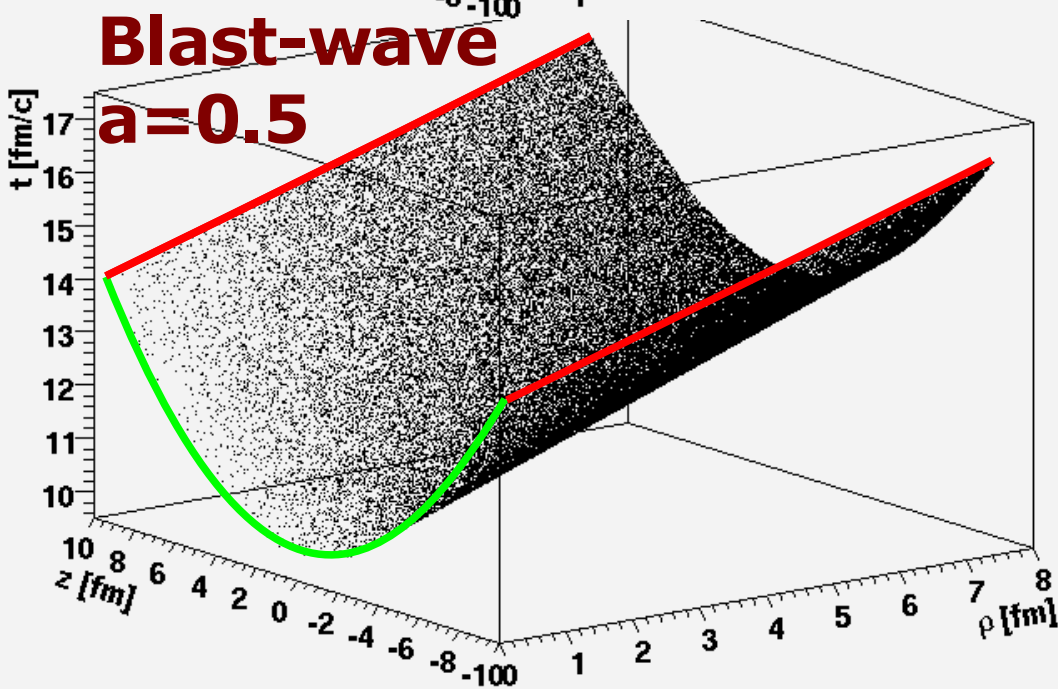
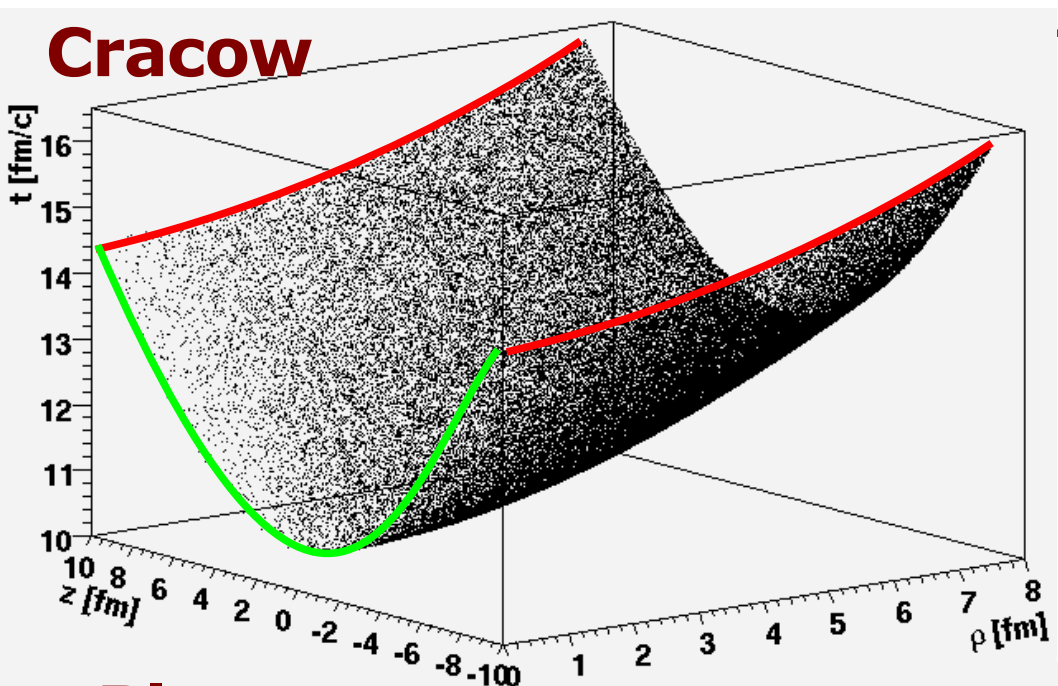


3D correlation functions with Coulomb

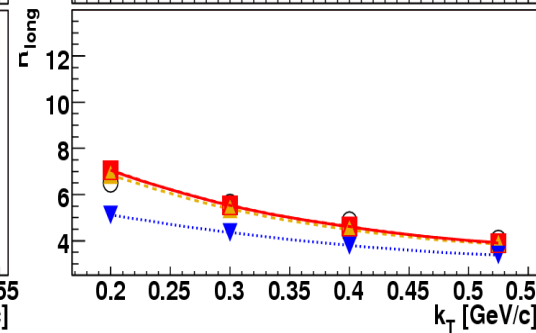
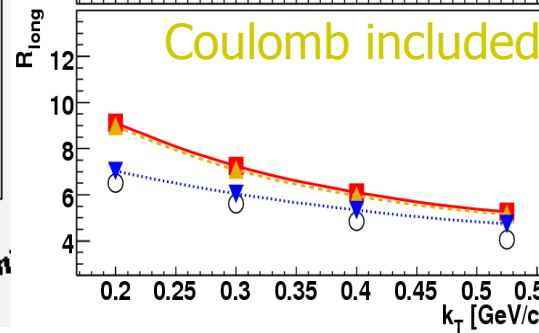
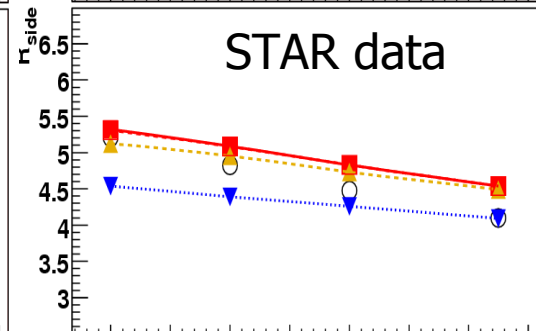
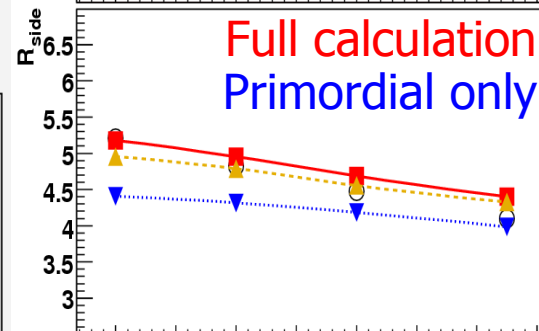
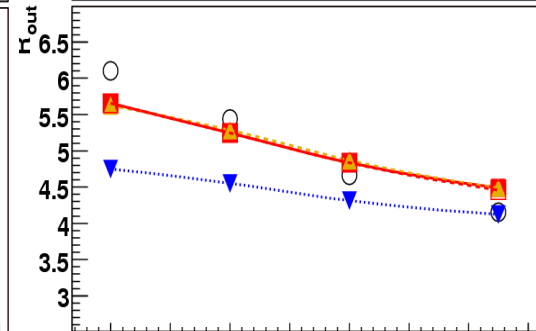
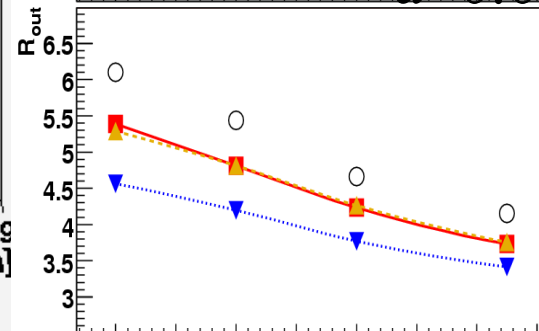
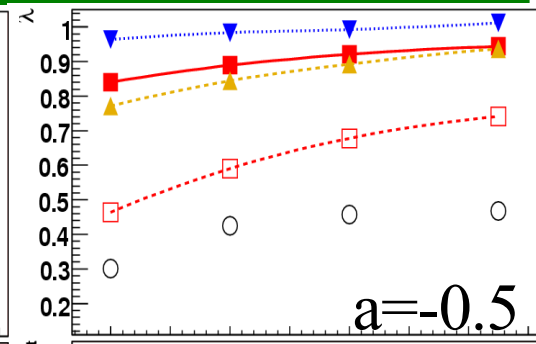
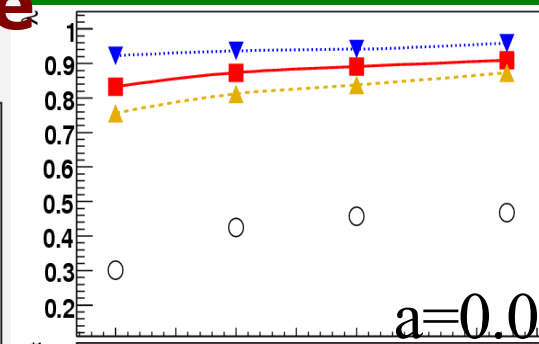
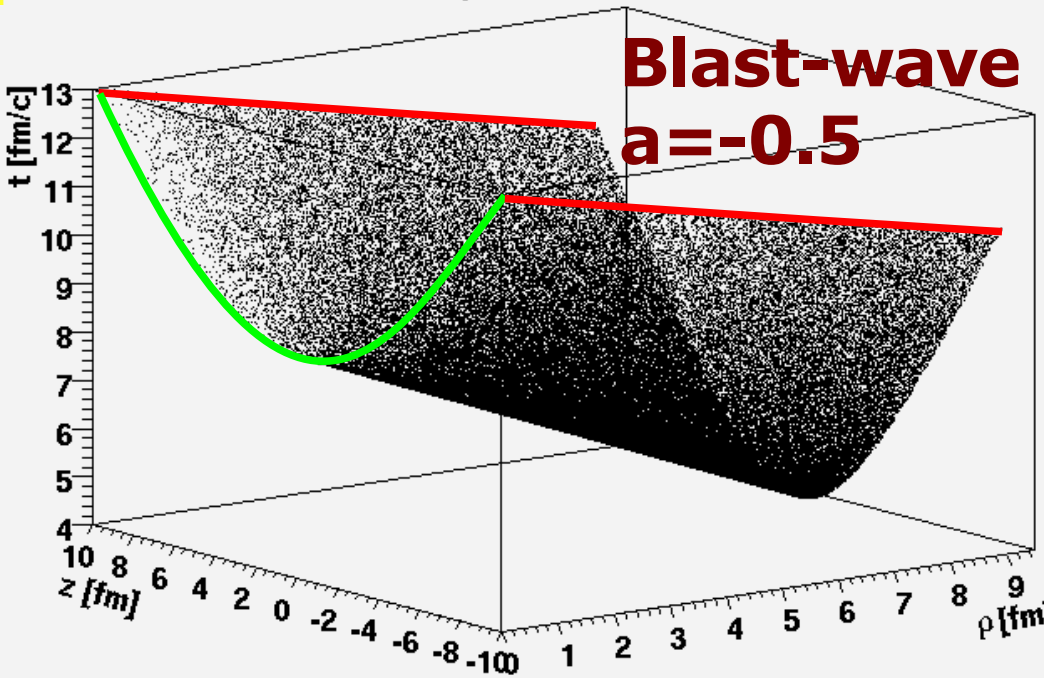
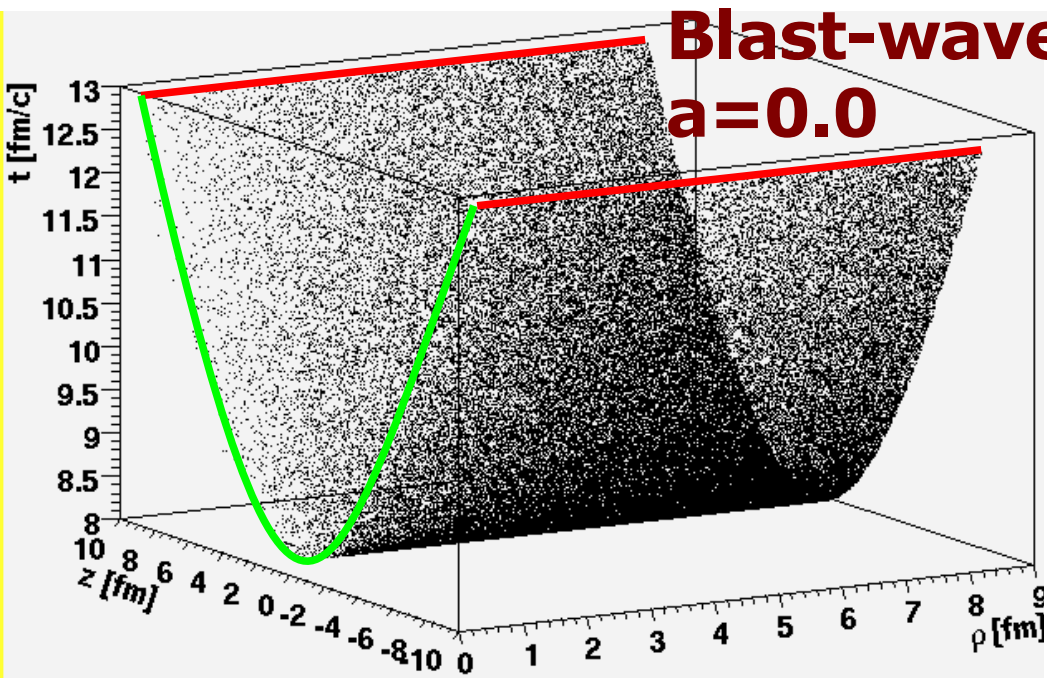


- The correlation function with full simulation of Coulomb effects can also be calculated
- It is fit with Bowler-Sinyukov formula. The fit is fully 3D. To plot it, we project it in the same manner as the input function.
- In this case we try to reproduce STAR data, therefore the K_{coul} for the spherical gaussian with radius 5fm in all directions was used.

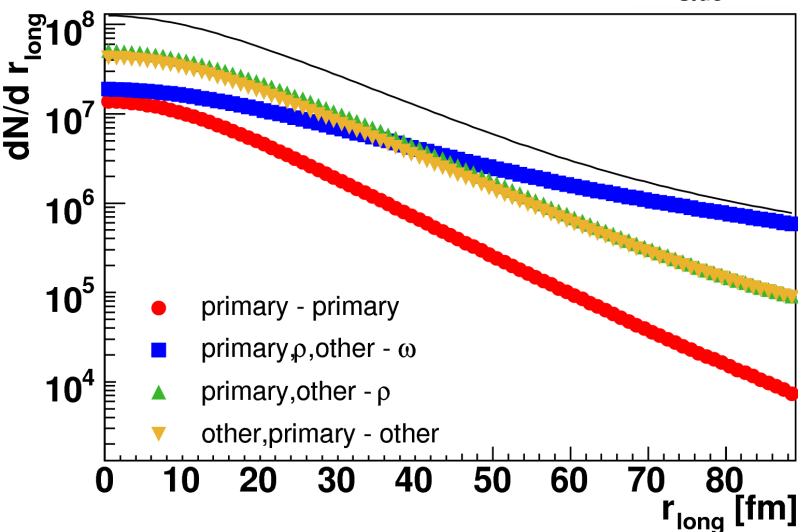
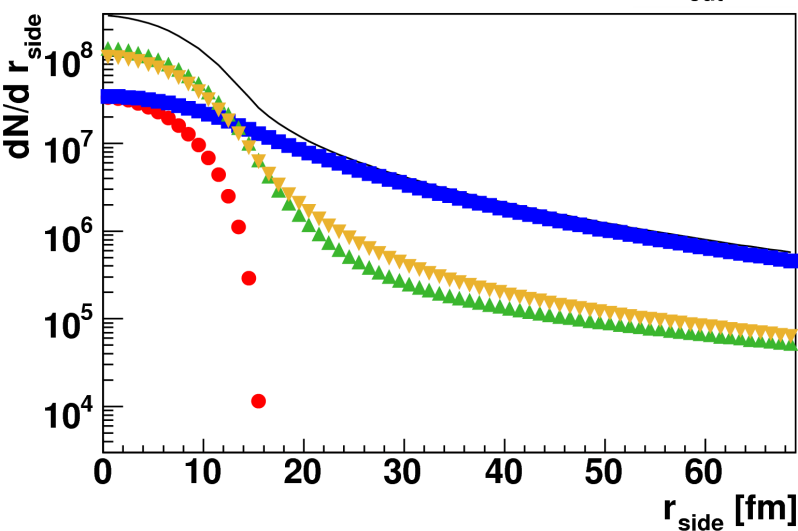
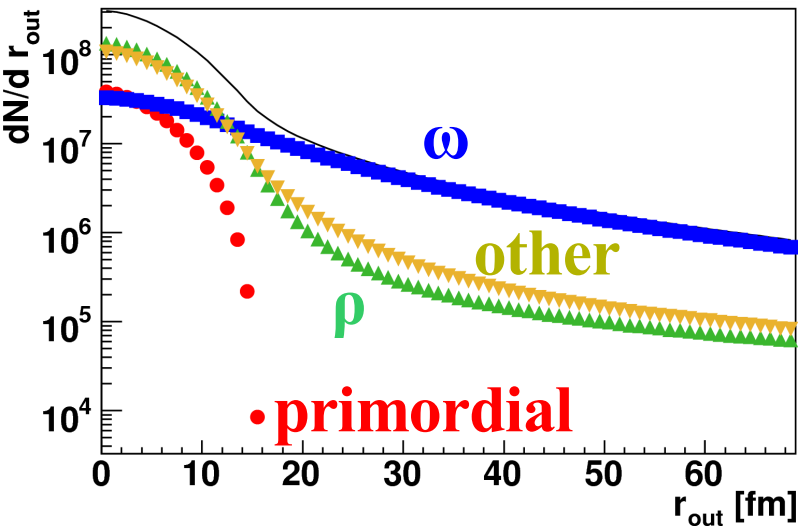
Comparing hypersurfaces



Modifying r-t slope

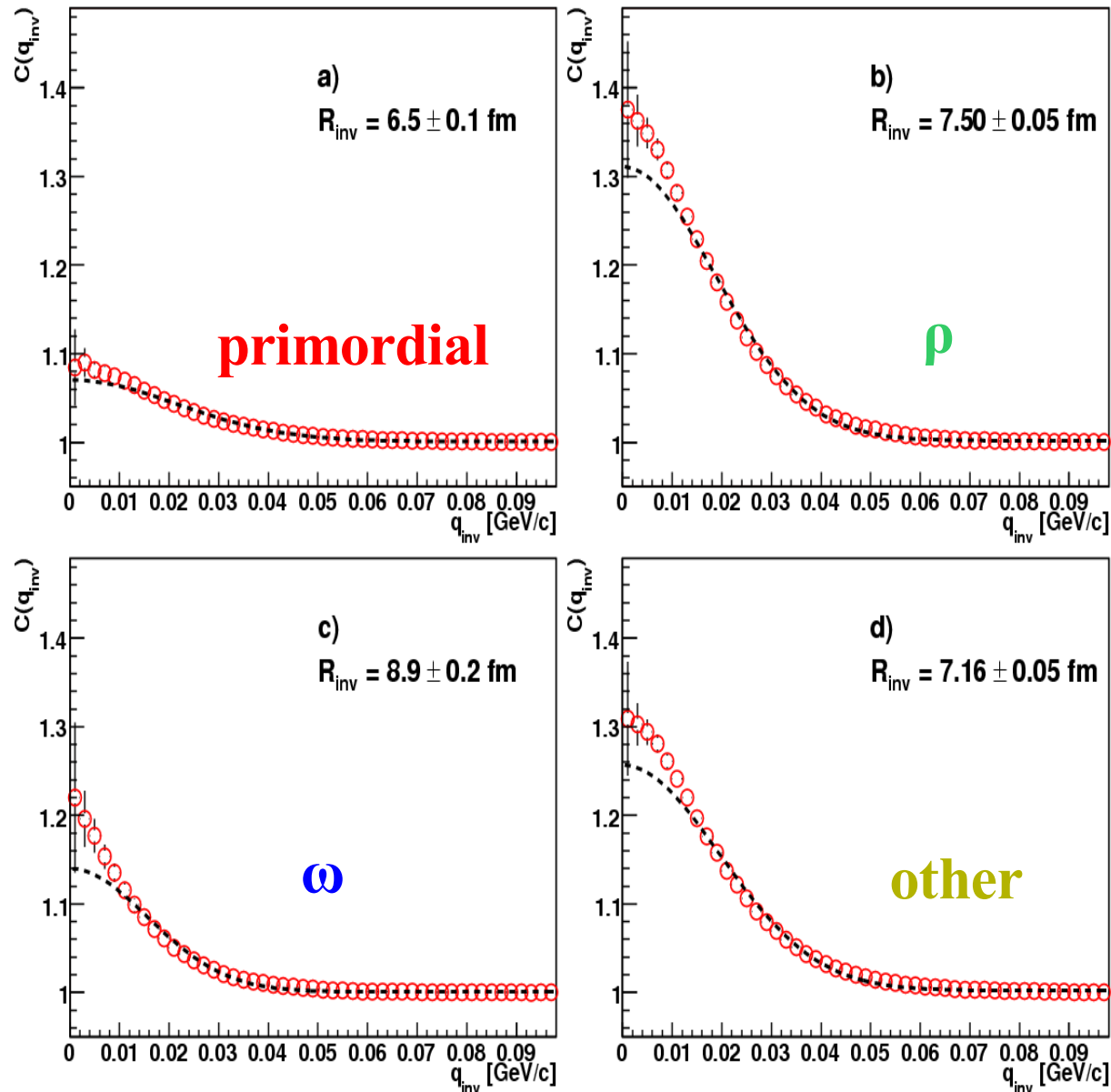


Resonance contribution



- Resonances have significant influence on the separation distributions – they produce long tails and enlarge the source
- The shape of the source is significantly non-gaussian
- The effect in the long direction is mixed with the influence of longitudinal expansion
- The source is completely non-gaussian in the long direction

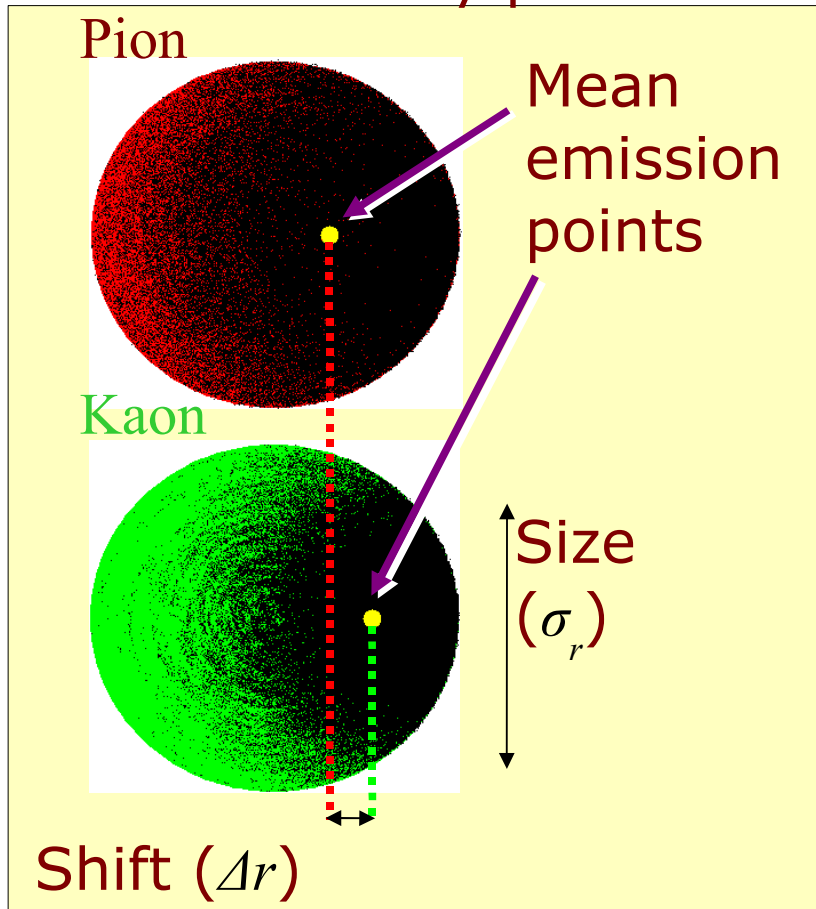
Influence on the correlation function



- Functions non-gaussian, but width roughly reproduced
- Primordial pairs give only 10% of correlation effect
- Resonances increase the size by about 1fm
- Contribution from omega sharply peaked – mostly visible in the lambda parameter

Femtoscscopy with Coulomb FSI

Pair velocity direction
Close velocity pair



- Hydrodynamic calculations with radial flow predict two effects:
 - Size decreases with particle m_T (length of homogeneity)
 - Mean emission point is shifted from the center

- Space effects are mixed with time shift in PRF:

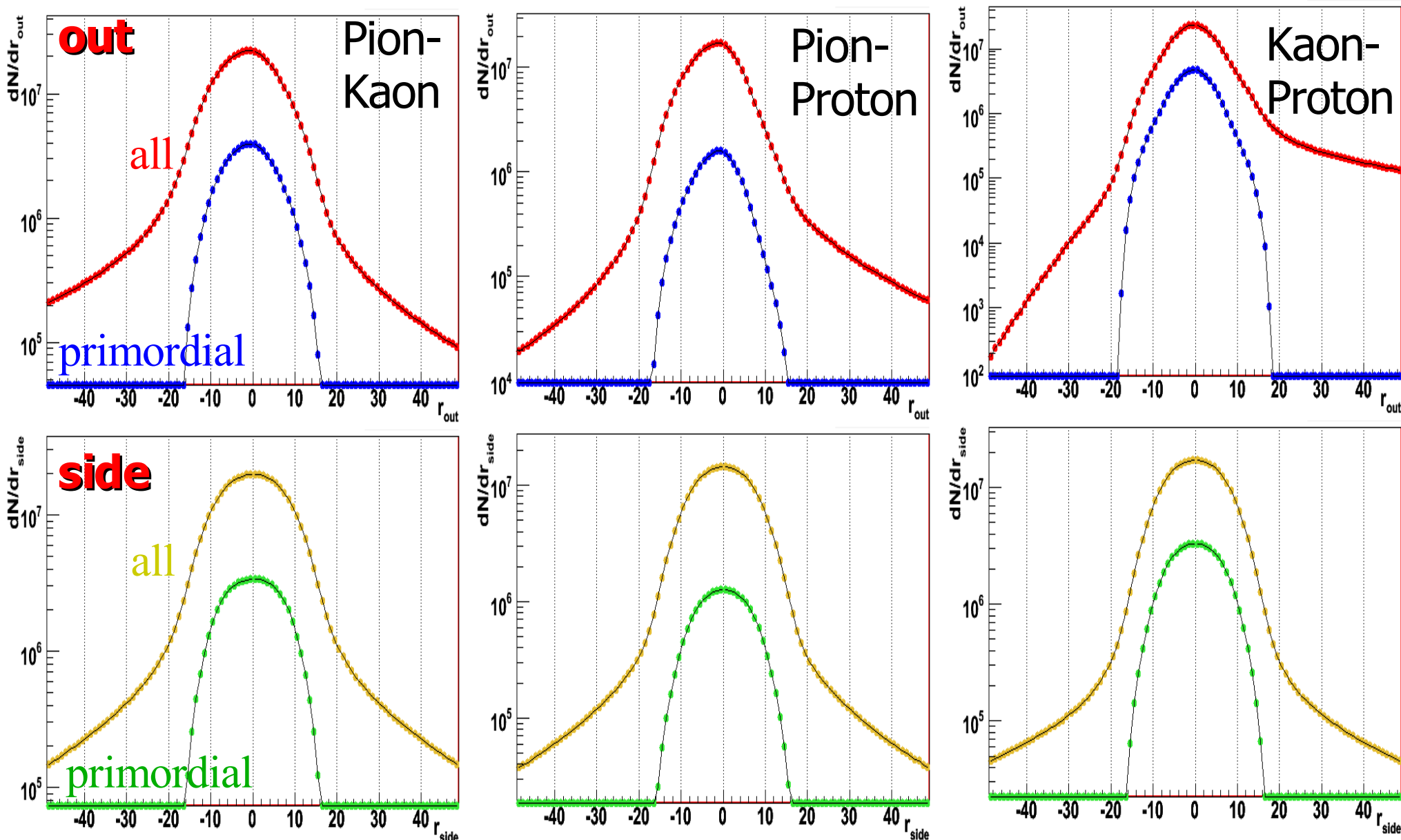
$$r_{out}^* = \gamma (\Delta r_{out} + \beta_T \Delta t)$$

it is crucial to simulate resonance contribution in detail

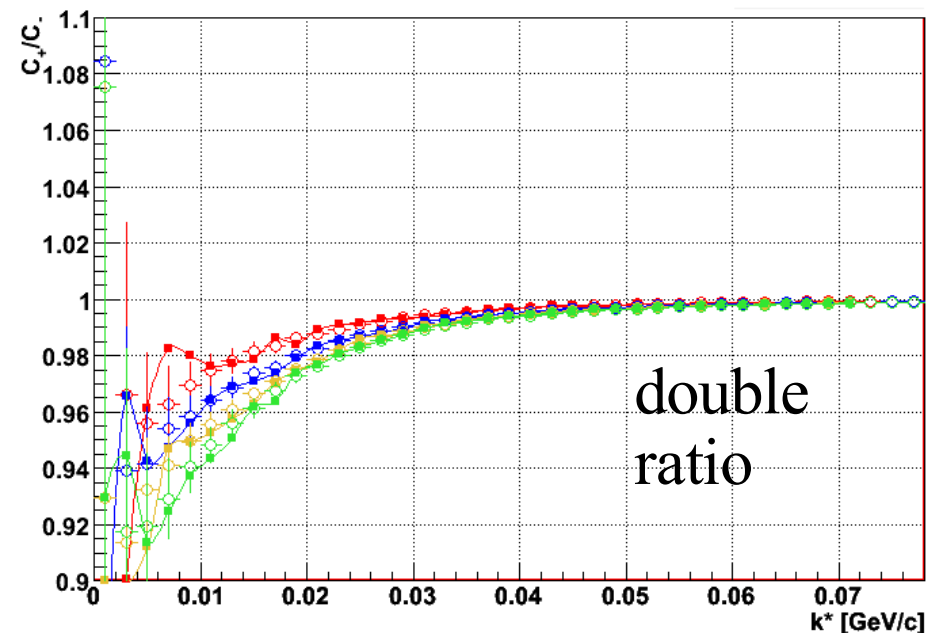
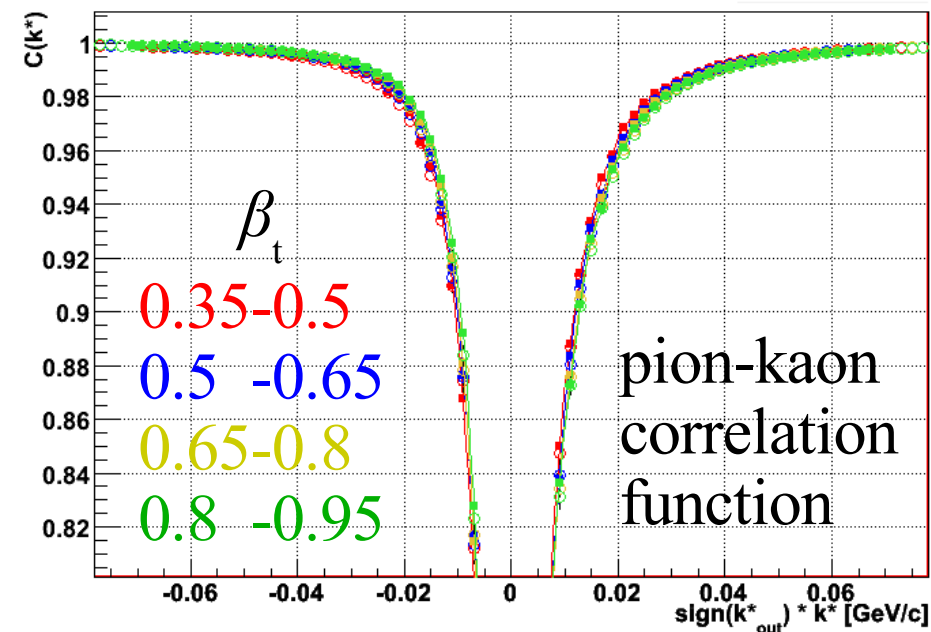
R.Lednicky et al. Phys.Lett. B373 (1996) 30.

S.Voloshin, R.Lednicky, S. Panitkin, N.Xu, Phys.Rev.Lett. **79**(1997)30

Separation distributions for non-identical systems



Non-identical correlation functions

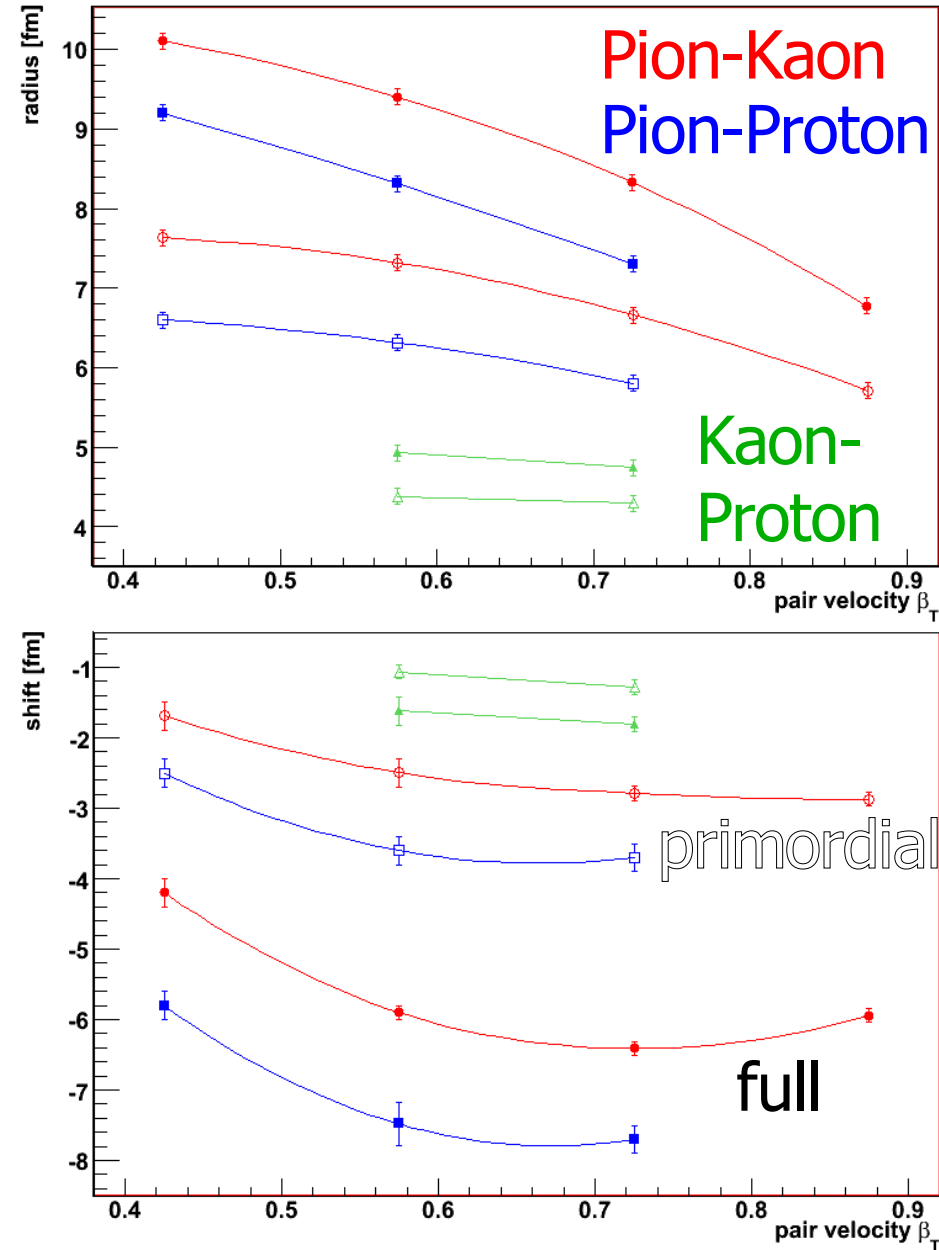


- We fit the model CF with a one simulated with the source function assumption:

$$S(\vec{r}, \vec{K}) \sim \exp\left(-\frac{(r_{out} - \mu_{out})^2}{\sigma^2} - \frac{r_{side}^2}{\sigma^2} - \frac{r_{long}^2}{\sigma^2}\right)$$

- Only 2 parameters (σ, μ_{out}) can be independent if we fit 1D function
- The correlation strength gives the size, the asymmetry determines the shift of average emission point

Non-identical particles results



- We observe strong asymmetries in pion systems
- Resonance propagation has large influence on asymmetry
- We observe the m_T scaling of sizes similar to identical particle interferometry
- Relative abundances of resonance to primordial particles important in quantitative analysis

Summary

- “HBT radii” are obtained from the blast-wave type of models with all resonance included, by performing full 3D simulation of the correlation function, with the possibility to include Coulomb effects
- Comparison to STAR data shows the single freeze-out model to be consistent with data, and preferring the freeze-out hypersurface with negative r-t slope.
- Resonances are found to influence all aspects of the correlation function: enlarge the size, decrease lambda, produce non-gaussian tails in the pair distributions
- First results on non-identical particle correlations study were presented